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THE SOLAR-POWERED SPACE SHIP

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Krafft A. Bhricke

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THE SOLAR - POWERED SPACE SHIP

by

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#### ABSTRACT

The characteristics of different propulsion systems for space vehicles are discussed and compared. The solar-powered space ship is evaluated in greater detail. A new light-weight design is presented, using spherical reflectors. Problems of design and operation are discussed by example of a hydrogen-operated prototype design. A number of basic characteristics of the vehicle system is established. A theor tical analysis of the spherical reflector as energy collector is presented.

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### Momenclature

- Area thrust gravitational acceleration Δħ enthalpy difference Specific impulse (1b thrust sec/lb fluid) Isp Q Total radiation energy flux density of reflector. flux density for unit minimum area (cf. appendix) R reflectivity at normal incidence R reflectivity radius of the spherical mirror solar constant ( 2 cal/cm<sup>2</sup> sec) temperature Stefan Boltsmann constant (0.56686 10-4 erg/cm<sup>2</sup> deg<sup>4</sup> sec 1.354410-12 cal (mean)/cm<sup>2</sup> (K sec)
  - For additional astations of. Fig. 11.

#### 1. Introduction

The enormous cost of supplying spaceborne vehicle systems from the earth is one of the principal restrictions in space flight with chemically powered propulsion systems. The vertetility and freedom of operations in space will increase to the extent to which the dependency on terrestrial supply of propulsion material can be reduced.

The dependency on terrestrial supply has two aspects, the first pertaining to the energy source, the second to the expendable metter. Thrust is produced by energising matter, theoretically even to the point of converting it into radiation In all other, less extreme cases, thrust is produced by accelerating expandable matter. As the exhaust velocity increases, the mass consumption decreases for a given operation; but the energy consumption of course increases. Thus, one faces the engineering problem of reducing the mass consumption (hence the terrestrial supply requirements) and increasing the energy supply and still maintaining reasonable overall operating conditions. This problem is not solved by simply pointing out another possibility for producing enormous exhaust velocities. one must also ask under what conditions of energy supply and thrust per unit weight these exhaust velocities can be obtained. In many instances a more detailed analysis shows that the atternment of such exhaust velocities is the least of the angineering problems involved and that the energy source, energy conversion etc. are now much bigger headaches; in other words, the problems are just shifted into another area, because one does not obtain anything for nothing. If one deals with such questions for a while, once comes to feel somewhat applopetic towards the "good old" chemical rocket which in many respects is indeed hard to best.

<sup>2)</sup> Mumbers in parenthesis refer to References on page

This is not intended to be a criticism of attempts to investigate other propulsion systems, but rather to caution lest one loses a realistic perspective towards the facts of space propulsion or feels that further hard work in improving the chemical system can be dispensed with in favor of more advanced systems. Space flight will become a reality through the chemical propulsion system. On the other hand, however, it is a fact that the supply requirements of the chemical system are a real handicap to anything more than occasional space expeditions. This shortcoming is significant enough to encourage research and analysis of alternate propulsion methods. The present paper attempts to make a contribution to these afforts.

#### 2. Propulsion System Characteristics

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With the possible exception of the hypothetical total photon propulsion, the main difference between chemical propulsion and all other methods of propulsion is the separation of energy source and working fluid, hereafter briefly called medium (Fig. 1). In the chemical system this consolidation of energy and medium provides smooth (comparatively) and rapid conversion, but limits the selection of media and maximises the dependency upon terrestrial supply, inasmich as oxidisers and fuels are nowhere readily available under natural conditions.

Separation of energy source and medium results in more flexible systems, but also raises new problems regarding energy supply, energy conversion and equipment weight. Fig. 1 presents a number of these systems which already have been suggested before on several occasions, e.g. (2) to (9), with the exception of the arc heating system which however has been studied previously for the purpose of producing an ultrahigh-speed gas flow in test facilities (10,11).

7

Briefly, the principal potential pros and cons of separation of energy source and medium can be summarized as follows:

#### Adventages:

- (a) Possibility of <u>higher</u> energy concentration than in chemical propellants through the use of nuclear power sources (e.g. pile,  $\beta$ -decay battery).
- (b) Possibility of permanent energy supply from the sum, at least in terrestrial and intra-terrestrial space. The energy supply, however, is much less concentrated then in either nuclear or chemical sources.
- (c) Greater freedom in the selection of media.
- (d) Possible simplifications of the propulsion system where only one type of fluid is used.
- (e) Possibility of supply of medium from other sources than the easth (e.g. refilling of the Saturn moon Titan or on Jupiter moons), thereby increasing the range of the ship for a given terrestrial supply.

#### Disadvantages:

- (f) Greater complexity and weight of the energy source, energy conversion and transfer mechanism and related equipment.
- (g) In many cases very low thrust-to-weight ratios, hence the danger of high gravitational losses when operating near planets as well as poor mancuverability. Application restricted to space only.
- (h) Difficulties in handling and maintenance of a nuclear energy source (pile).
- (i) Operational difficulties resulting from excessive energy release necessitated by low conversion efficiencies, such as the need to dispense with excess heat in conversion systems.

The last mentioned disadvantage becomes the more severe the higher the ultirate energy level prior to conversion is supposed to be, that is, the more energy is available prior to expansion and/or discharge of the medium. Figure 2 presents a survey of the energy converted in producing thrust versus the specific impulse for various propulsion systems. For reasons of comparison the equivalent energy for different flight mechanical energy levels is given. The relatively flat slope of the conversion line shows that any increase in specific impulse must be bought at considerable increase in energy imparted to the jet (in fact,  $I_{an} \propto \sqrt{\Delta h}$  ). This is the reason for the problem shift towards the energy source and conversion system as the specific impulse goes up; and obviously, if 10 kcal/g instead of 1 koal/g is required at 0.25 overall conversion efficiency, the production of 40 koal/g at the source and the need to dispense with 30 kcal/g is much more of a problem then the production of 5 kcul/g and the need to remove 3 kcal/g. Thus, high specific impulse is desirable, but it must be in proportion to the flight mechanical energy requirements (i.e. not greater than necessary and convenient for the contemplated mission).

A survey of some basic characteristics of these propulsion systems is presented in Tab.1. Most items are self-explanatory. It should be pointed out that the specific energy consumption in kw per 1b thrust was evaluated as follows:

The production in the propulsion system to get the thrust producing mechanism started, the energy converted from its original form into thermal energy, and finally the energy converted in the jet proper, which is equal to Δh(Btu/sec)/0.984 I<sub>sp</sub> (1b sec/lb), where Δh and I<sub>sp</sub> are taken from Fig. 2 and 1/0.984 is the conversion factor from Btu to Kv. The figures pertaining to the chemical system reflect the fact that little has to be produced outside the medium (only auxiliary power) while the medium itself releases the energy. The conditions are reversed in all other cases where energy must be transferred to the jet.

Against this general background the solar-powered space ship will be discussed in more detail.

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Table 1 shows that among the possible non-chemical systems, the solarpowered drive is fundamentally the simplest and most straight forward arrangement. It uses an existing energy source in space in an efficient manner, by converting the radiant energy directly into heat. The propulsion system simply consists of tankage for one type of fluid, radiation collector, heat exchanger and exhaust nossle. Fumps are required to circulate the working fluid, but since the quantities involved are very small (less than a pound), the pumps are light, the horsepower requirement is low and electrical high-speed drives appear practical. The necessity of concentrating the thin-spread solar energy requires the use of large reflectortype collectors. In this manner very high heat flux densities can be obtained which are equal and greater than those normally found in the throat of chemical rocket engines (> 3 Btu/in sec). Most fluids are heated to a very high temperature under these conditions and this makes their subsequent piping to the exhaust nossle quite difficult. If hydrogen is selected, the high heat capacity permits to store a considerable amount of energy in the fluid, while keeping the temperature within such limits (1,500~1,500°F) as to permit the use of uncooled pipelines for the heated material, while the specific impulse nevertheless is the highest attainable with any fluid under these conditions and, in fact, exceeds that of chemical propellants (Figures 3,4). By using uncooled pipelines, the system can be simplified and its weight kept low.

Low weight is of extreme importance, since the thrust obtainable is only of the order of 100 to 200 lb. Even in the case of orbiting space ships the thrust-to-weight ratio cannot be allowed to become arbitrarily low. The lower limit is rather determined by the specific impulse available and by the strength of the gravitational field in which the system is to operate. In the case of the earthmoon field and with a specific impulse around 450 sec, the thrust-to-weight ratio should not rall appreciably below 0.01g, because otherwise the flight mechanical performance during the powered phase becomes so poor that impractically large mass

ratics are needed to perform a cialmar flight mission. This is due to the gravitational losses incurred in a slow climbing spiral-path associated with very low thrust. However, low thrust is not only a disadvantage. It is in fact mandatory for the system under consideration, because the considerable size of the collectors yields very long moment arms. At thrust-to-weight ratios of 0.1 to 0.7 g it would not be possible, within reasonable weight limits for space vehicles of this size, to provide the macessary structurally rigidity and to prevent lending and distortion which would destroy the optical quality of the reflectors.

A compromise between these two opposite thrust requirements must be established and leads to values of the order of 0.01 g. If the right conditions are fulfilled, the solar-powered drive will require much less working fluid supply from the earth than any chamical propulsion system. The limited power of this system takes its use for interplanetary flights unlikely. It appears suitable for cichaar and lunar operations which presumably will be more frequent and for which a reduction in fluid supply is therefore quibe important.

However, there are very problems to be solved if its usefulness is to be assured. The principal design problem lies in the extreme emphasis for all-out light-weight construction. Another design as well as dynamics problem is the requirement for completely independent orientation in space of the optical axis and the thrust axis with respect to each other. The problem of rigidity and the resulting autopilot control difficulties can be reduced to practical values by reams of proper bracing of the collectors. Considerable pressure leases are introduced by long pipelines between tank, heater and motor, requi ing additional pumping energy. This difficulty is apparently unavoidable. A large number of different designs has been evaluated with the purpose of reducing the length of piping, before the design presented in this paper was selected. Basic shortcomings of the system are of course dependency on solar radiation and the possibility of damage even by small meteors due to large size and freil construction. The vehicle is

without means of propulsion while in the shadow of a celestial body. The effect of cosmic dust may be the most serious of all problems and may in fact constitute a hurdle which cannot be overcome. It is too early, however, to assess the effect of dust in space with sufficient certainty to make such a statement. Furthermore, the density and distribution in cislumar space must be known more accurately. It is not to be expected that the distribution will be uniform in view of the complex interaction of terrestrial, lumar and solar gravity fields in cislumar space.

Further technical problems will become apparent during the subsequent discussion. Altegether they make the solar-powered space ship difficult to realize; however, while constituting a great challenge to engineering ingenuity, the successful solution of its problems will reward us with increased freedom in space.

#### 4. The Working Fluid

The selection of the working fluid is mainly determined by the requirement for low overall vehicle weight and the definite need for high specific impulse. Desirable qualifications derived from the preceding discussion are high heat capacity and low condensation temperature to permit a maximum degree of expansion and high energy conversion at relatively low initial temperature. In order to maintain a high heat transfer coefficient throughout the heating period, the working fluid must be in supercritical state. Therefore a not too high critical pressure is desired.

The working fluids which best meet these specifications are hydrogen and helium. Their relevant characteristics are summarised in Table 2. Hydrogen yields higher specific impulse, because its molecular weight is lower and its specific heat higher. Helium has the advantage of higher density and of lower critical pressure as well as higher critical density. It has also a lower heat of vaporisation. The density impulse of helium is about 10 percent higher than that of hydrogen.

However, it is believed that hydrogen is in this case the only choice which is acceptable from practical considerations. It is more readily available, it permits to operate the probalision system at a much lower temperature level at about equal motor performance and it yields a lower gross weight, hence a higher thrust-to-weight ratio and less gravitational losses. At present, sufficient practical experience in the use and pumping of hydrogen in rocket engine systems is available to permit estimates for the technical layout of a solar-powered propulsion systems. The molar heat of hydrogen under different supercritical pressures is shown in Fig.5. These curves were obtained by measuring the slope  $(dS/dT)_p$  of curves in the hydrogen entropy-temperature diagram ref. (14) and computing the molar heat capacity from  $G_p = (dS/dT) \cdot T$ .

An estimate of the density of laquid hydrogen under the slightly supercritical pressure of 15 atm is shown in Fig. 6. The variation of exhaust temperature  $T_e$  and enthalphy  $\Delta h$  converted in the nozzle is shown in Fig. 7 as function of chamber pressure and chamber temperature. Figure 7 also shows the enthalpy  $\Delta h$  required to convert the hydrogen from its original state in the tank to the state prior to expansion in the nozzle. Because of the relatively low temperatures involved, dissociation is nebligible and  $\Delta h$  is not a function of the pressure prior to expansion. By dividing  $(\Delta h/\Delta h^*)_T$  one obtains ratios between 0.75 and 0.85 at supercritical pressures.

#### 5. The Energy Collector

In order to increase the enthalpy of the working fluid by the required amount in a short period of time, the radiation energy must be concentrated in a small area in which the heater is placed. Usually, parabolic reflectors are considered for this purpose (4, 7). In a parabolic reflector the reflector is concentrated in a small focal area where extremely high temperatures or heat transfer rates can be obtained. Farber and Davis (16) have analyzed the parabolic reflector for the purpose of attaining maximum temperatures in the focus. They find the highest theoretically possible temperature of a black body receiver to be 5,100°K (assuming 100 % reflectivity), compared to 6,000°K solar surface temperature.

In a solar propulsion system the use of a collector is not to attain highest heat transfer rates in the smallest area. The purpose of the collector is to distribute energy over a certain area so that sufficient time is given the fluid to absorb energy while at the same time the flux density remains high enough to produce a source-sink system of adequate intensity for the temperature to be attained by the fluid. In order to do this with a parabolic reflector, its optical quality must be reduced purposely.

However, the main argument against the use of a parabolic mirror comes from weight considerations in relation to the energy obtained or thrust generated per unit of area intercepted. The radiation energy collected by a reflector is equal to the area intercepted, A, times the selar constant, S, and the reflectivity at normal incidence,  $R_n$ ,

 $Q = A \otimes R_n \tag{1}$ 

From Fig. 7,4h\* for  $T_c = 1,000^{\circ}$ K, neglecting losses, is 3,557.5 calories per gram of hydrogen. The solar constant in space is S = 7/30 cal/sec cm<sup>2</sup>. Thus, if

one assumes roughly that all energy intercepted can be transferred to the working fluid (hydrogen), one obtains for the reflector area per unit weight of hydrogen per second, heated to 1,000°K,

$$\frac{A}{(\hat{\mathbf{v}}_{H_2})} = \frac{106,725}{R_n} \frac{\text{cal sec cm}^2 - \text{cm}^2}{\text{g sec cal}} \frac{\text{cm}^2}{\text{g }(H_2)}$$

$$= \frac{10,672.5}{R_n} \frac{\text{m}^2}{\text{kg}(H_2)}$$

$$= \frac{52,190}{R_n} \frac{\text{ft}^2}{\text{lb}(H_2)}$$
(2)

From Figure 3 one obtains for a pressure ratio of 15/0.1 atm through the nozzle and  $T_{\rm c}=1,000^{\rm o}{\rm K}$ , a specific impulse of 478 sec. . Assuming 94 percent or 450 sec. one obtains for the reflector specific impulse

$$\frac{A}{I_{sn}} = \frac{116}{R_n} \frac{ft^2}{lb(thrust)} = \frac{23.7}{R_n} \frac{m^2}{kg(thrust)}$$
(3)

Since the energy intercepted is about 1 kw/m<sup>2</sup>, one arrives at 10.8 kw/lb(thrust) at  $R_n = 1.0$  or 12 kw/lb(thrust) at  $R_n = 0.9$ . The additional weights of propulsion, structure, such as gondola etc., and hydrogen itself leave very little weight for the collector system. In order to arrive at an overall thrust-to-weight ratio of at least  $10^{-2}$ , the collector weight (lb) to thrust (lb) must be about 5;1 to 7:1. This then allows about 0.043 to 0.06 lb weight per ft<sup>2</sup> intercepted area for the collector system. Such values can obviously not be realized with a parabolic reflector for structural reasons.

During discussions of this problem with the author's associates

Mesors. F. D'Vincent, C. Edenfield and O. Dahlke<sup>3)</sup> the proposal was advanced
to use a thin-walled, presoure-stabilized sphere. Such a system not only has
the least possible weight, but also yields a naturally correct reflector which

Senior Design Engineers, Pre. Des. & Syst. Anal., Convair, San Diego. The author is deeply injebted for their help without which the design could not have been advanced to the level presented here.

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in this case would be a hemispherical reflector. The reflector would be produced by spraying one half of the sphere with a thin metallic film of aluminum or silver, while the other half would remain highly transparent. The resulting weight reduction is so large that a certain increase in intercepted area to compensate for slight absorption losses on the transparent side can easily be accomplated. An analysis of the hemispherical reflector is presented in the Appendix and it is shown that heat transfer rates of maximum practical intensity are obtained along that part of the optical axis which represents the focal line in a reflector of this shape. It was therefore decided to investigate a vehicle design on this basis.

#### 6. The Solar-Powered Space Ship Prototype

For this purpose, a small hydrogen-operated prototype of about 16,000 lb gross weight was assumed, containing about 11,000 liquid hydrogen and a gondola for two crew members. Several configurations were considered. The eventual design which is shown in principle in Figure 8 and in more detail in Figure 9 was chosen, because it appears to be the most attractive compromise from a number of viewpoints.

Gondola, spherical hydrogen tank and exhaust nozzle represent one rigid system, located in the center between two collectors. Center portion as well as reflectors can be rotated freely with respect to each other about the axis of rotation connecting the centers of the two collector spheres. Large vanes in the motor exhaust provide roll control (about the thrust axis). By means of combined motion about axis of rotation and thrust axis, as well as yaw control about the third axis normal to the plane of the paper, complete freedom of orientation of thrust axis and optical axis with respect to each other is assured. Yaw control is effected by means of tilting the exhaust nozzle, using actuating cylinders which also provide a possibility for fine control of the thrust about the axis of rotation by having these actuators operate in two different planes, as explained in Fig. 9, Detail A. Detail B explains the connection between collector sphere and tank sphere and shows that an electric drive is provided for each collector to rotate with respect to the thrust axis. The thrust axis in turn can retate with respect to the collectors by tilting the motor normal to the plane of the paper. The vehicle can thus accelerate in any direction with respect to the sum. The arrangement of equipment in and around the collector spheres is such that the center of gravity of each collector is in the center of the sphere, or at least on the axis of rotation, so that the reflectors can be rotated without producing a moment arm with respect to the line of thrust.

The flow diagram Fig. 8 shows that hydrogen is pumped from the tank into the collector sphere. This is done by means of a low-pressure pump to keep the flow velocity and therewith the pressure losses down during the about 74 ft long flow to the center of the sphere. If necessary, additional pumps must be installed along the way to maintain the flow as the desired booster pressure prior to entering the high-pressure pump in the center of the collector. Since the flow quantities are small - of the order of 0.1 to 0.2 1b for each collector - and since the pressure is relatively low, the weight of these booster rumps together with their electric drive is very small, of the order of a few pounds. In the high-pressure pump system, the pressure is increased to the order of 30 atm (440 psi), to intensify the heat transfer. Subsequent pressure losses on the way to the motor are not allowed to reduce the pressure below 15 atm, using again booster pumps if necessary. Expansion through the nozzle takes place at an initial pressure of about 15 atm (220 psi). The arrangement of the high-pressure pump close to the heating element reduces the length of high pressure lines. Assuming that the inlet pressure at the high-pressure pump is 30 psi and the outlet pressure 450 psi and assuming further an efficiency of 0.5 for the pump and 0.8 for the electric motor, the horsepower requirement for each of the two systems is about 9 HP or 6.7 kw. This and additional nover is probably most economically provided by means of a turbo-electric system. In Fig. 9 a single-stage impulse turbine is indicated prior to the expansion nossle. Py means of an alternator AC is produced ( 3 phase, 4 wires, 120/208 V) which has the advantage of yielding lower transmission losses and permitting a higher rotational speed of the motors and pumps than with DC where the motor speed is limited to about 8,000 to 10,000 rpm. Of the above mentioned 4 wires, 2 each lead to the collector spheres. By keeping them in close contact with the cold hydrogen lines, transmission lessen can be minimised. These losses are estimated to be about 0.25 HP per leg.

For starting the propulsion system a high-pressure gas tank for the initial pump operation is indicated in Fig. 9. However, for repeated starts solid propellant starter rockets to energize the turbine and produce a small amount of thrust appear preferable.

Energy for auxiliary power needed in the gondola and for radio purposes can be made available either from radiation of lesser intensity in the heater area or from special solar batteries shown at the extension of the respective optical axes of the collectors in Fig. 9. Storage batteries take over while the vehicle passes through the shadow of the earth. The solar batteries outside the collector spheres would block out radiation arriving in a cylindrical space around the optical axis. It is shown in the Appendix that this radiation furnishes only a negligible contribution to the energy concentration along the optical axis (focal line).

The heater extends from a point half way from the center to a point on the periphery of the reflector. Figure 10 shows the intense heat flux density along the focal line for several values of the reflectivity (cf. Appendix). Peak values between 8 and 10 Ptu/in<sup>2</sup> sec are reached at  $\varphi$  about 45 degrees, corresponding to approximately 70 percent of the distance from the center to the periphery (Appendix).

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The spherical collector proper is assumed to consist of polyester (polyethylene terephthalate), a clear and transparent plastic of considerable strength and very light weight. Very thin films of polyester can be manufactured and the design shown in Fig. 9 is based on a thickness of 0.001 inch. Under these conditions the material is not expected to absorb any appreciable amount of light or to produce significant refraction. In order to utilize the industrial state of the art, a polyester called Mylar D which is being produced, has been evaluated (17). The subsequent information, representing average values, is drawn from this reference and is summarized in Table 3. The polyester film transmits about 90

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**(**)

percent of the incident light in the visible region. The tensile strength is satisfactory, even at 300°F (150°C). Under the given conditions of thrust-to-weight it appears sufficient to pressurise the sphere with hydrogen or helium at 6.01 psi (about 7.10-4 atm) to lend adequate rigidity to the hemispherical reflector. This yields a skin stress of 3.840 psi which is well under the tensile strength left at 300°F. This is approximately the daylight temperature of the lunar surface. It is expected that the film temperature will stay below this level, since it absorbs less radiation than the moon's surface. Moreover, since hydrogen or helium will diffuse through the skin, especially on the transparent side, a certain amount of cooling is provided automatically where it is most desirable. The very low pressure on the other hand keeps the amount of gas pressure to be renewed on a very low level. For the same reason also micron-size holes punctured into the film by cosmic dust are not expected to be critical, depending, of course on the density of the dust. To a certain extent the transparent film and the internal gas which in spite of its rarefaction contains more than 1020 molecules per ft3, provide protection for the reflector proper against frontal impingement by cosmic dust. The rear side is more resistant, since this side is expected to be metal-surayed on both, the inner and the outer side with a layer about one micron thick. The weight of the metallic film is negligible and the film thickness could be increased if desirable for reasons of protection from cosmic dust. In the direction of low temperature the polyester material shows also good qualities. Mylar is quoted in ref. (17) to be free from embrittlement at temperatures as low as - 60°C (-78°E). The thermal radiation given off by the heater when cooling down and the heat content of the thin internal atmosphere, will greatly dampen the temperature drop normally encountered when the vehicle enters the shadow of the earth.

Thus the material seems to be applicable to this design. However, it is realized that there are many unknowns left, particularly with respect to cosmis dust and the effect of exposure of this material to those parts of the solar spectrum which are absorbed by the atmosphere. However, it appears still too early

to rule out this type of application. Moreover, further directed development may lead to additional improvements. The problems encountered in this design are characteristic problems of an advanced space technology and must be solved jointly by industrial and satellite research.

A weight summary of the prototype is presented in Table 4. Both sollectors together weigh only 740 lb or, with accessories, 1,000 lb. This is roughly the equipment which replaces the 44,000 lb oxygen otherwise needed to heat the hydrogen (11,000 lb) chemically. It is this fundamental advantage which makes the solar-powered space ship a significant possibility and well worthwhile the effort to solve its numerous existing problems. Of course, this advantage is not all gain. There are certain penalties in the form of design and operational difficulties. Another, most significant penalty, namely the flight mechanical performance loss due to the low thrust-to-weight ratio, will be discussed in the subsequent section.

#### 7. Somery and Conclusions

The basic characteristics of non-charactal propulsion satisfies are surveyed. The principal reason for a change to these systems is the attempt to get away from the high mase consumption which represents the sajor disclosurage of the characteristics method. However, it is pointed out that in doing so, a host of other, partly even more severe difficulties is encountered. All non-chamical methods require densiderable research and development effect before they can become practical.

Because of its advantages in simplicity and high energy conversion efficiency the solar-powered space ship is discussed in more detail. This vehicle is potentially capable of carrying out operations in dislumer and lumer apage (without landing) on a much more economical basis as far as supply requirements are concerned. However, the system is a good example for the manifold difficulties encountered when changing to another propulsion system. These difficulties are not necessarily impresuntable, but their severity is considerable and the resulting problems present many challenges to engineering and science. In the case of the solar-powered space ship the requirement for extremely light construction is the determining factor, to a greater extent than ever before in the history of recket development which is the story of man's flight against weight in more than one sense. It appears that this problem can be overcome by radically new designs which must be based on the bast materials industry can provide. The prototype presented here may indicate a possible solution. It appears likely that if at a leter time industrial and scientific ingenuity can be put to work on a proader basis, s practical solar-powered space ship can be developed and man's freedom of operation in the earth-moon field be increased deci ively.

#### Appendix ; Analysis of the Spherical Reflector

Figure // defines the notation used subsequently. The optical or principal axis passes through the vertex V and the center 0 of the sphere. For the sake of simplicity of the geometrical relations involved, solar radiation is in this connection assumed to consist of parallel rays of light. The reflection point R of an incoming light ray can be described by the center angle  $\varphi$ , measured from a line normal to the incoming light and into the direction of the optical axis toward the vertex V. It is then

$$\frac{h}{r} = \cos \varphi$$

$$\frac{h}{f} = \sin 2 i$$

$$\frac{f}{f} = \frac{\cos \varphi}{\sin 2 \varphi}$$
(1-2)

and

The last relation defines the length f of the reflected beam in terms of the radius of the spherical reflector. This length, times the angles subtended by the light source, determine the diameter of the image. In the case of the sum this angle is known to be  $32^{\circ} = 0.00931$  radians =  $\beta$ . The smallest possible image diameter is therefore  $f \beta$  at any station of the optical axis irradiated, and it lies in the plane normal to the direction of f. If the element lies in the axis 70, the incident radiation will spread to form an ellipse of light rather than a circle. The major axis of this ellipse is given by

$$2a = \frac{f\beta}{\cos^2(90-i)} = \frac{f\beta}{\sin^2\varphi} \tag{1-3}$$

The radiation intercepted by the minimum area  $A_{\min} = (\pi/4)f^2/t^2$ , as compared to the total elliptic area irradiated is,

Elliptic area 
$$A_{ell} = ab = \pi \frac{1}{2} \frac{f\beta}{\sin 2\varphi} \cdot \frac{1}{2} f\beta = \frac{\pi}{4} \frac{f^2\beta^2}{\sin 2\varphi}$$

$$\frac{A_{min}}{A_{ell}} = \sin 2\varphi \qquad (1-4)$$

Assuming a flat plate receiver, the radiation absorbed is (for  $\alpha = \epsilon$ )

$$\alpha = \epsilon = \epsilon_n \cos(90 - il) = \epsilon_n \sin 2i$$
 (1-5)

For a small reflector surface element the incident flux is then given by

$$q = S \cos i = S \sin \varphi \qquad (1-6)$$

where S is the solar constant. This reduction in intensity is of course due to the spirading of the incident flux over a larger area than in the case of vertical incidence (at point V). The reflectivity of the surface element follows from Kirchhoff's law (assuming  $\alpha = \ell$ ),  $2\ell + R = 1$ ;  $\ell = (1-R)/2$  and Lambert's law  $\ell = \ell_R \cos i = \ell_R \sin \ell$ , whence

$$R = 1 - (1 - R_n) \sin \varphi \tag{1-7}$$

The amount of radiation absorbed by a flat plate receiver (diameter f 3 ) from a surface element af arc length ds, revolved about the optical axis, is therefore

$$dq = S \sin \varphi R \frac{Amin}{A_{ell}} \propto 2\pi (2h) f d \varphi$$
 (1-8)

where S  $\sin \varphi$  is the incident flux, R the reflectivity,  $A_{\min}/A_{ell}$  the fraction of radiation intercepted by the minimum area,  $\infty$  the absorptivity.  $2\pi(2h)$  the circumference at the given h about the optical axis and f d $\varphi$  the arc element. Using Eqs. (1-7), (1-4), (1-5) and (1-1) and setting f d $\varphi$  = ds, one obtains the flux density wer area element

$$\frac{dq}{ds} = 4\pi r 5 \epsilon_n f_i(\varphi) \left[ 1 - (1 - R_n) \sin \varphi \right]$$
 (1-9)

where

1

$$f_1(\varphi) = \sin^2 \varphi \sin \varphi \cos \varphi \tag{1-10}$$

Integration of this equation yields the overall flux from the reflector surface on the invadiated portion of the optical axis.

$$Q = 8 r^{2} s = n \left[ \frac{1}{5} \left( sin^{3} \rho cos^{2} \phi + \frac{2}{3} sin^{3} \phi \right) - \frac{1 - Rn}{6} \left( sin^{4} \rho cos^{2} \phi + \frac{1}{2} sin^{4} \phi \right) \right]$$
 (1-11)

The flux density in terms of the minimum area Amin is at any point

$$q = \frac{Q}{\frac{\pi}{4} f^2 \beta^2} = \frac{32}{\frac{f^2}{\gamma^2} \beta^2} \left[ f_2(\varphi) - \frac{(1 - R_n)}{6} f_3(\varphi) \right] (1-12)$$

where

$$f_2(\varphi) = \frac{1}{5} \left( \sin^3 \varphi \cos^2 \varphi + \frac{2}{3} \sin^3 \varphi \right)$$
 (1-13)

$$f_3(\varphi) = \sin^4 \varphi \cos^2 \varphi + \frac{1}{2} \sin^4 \varphi$$
 (1-14)

Eqs. (1-2), (1-4), (1-10),(1-13) and (1-14) are tabulated in Table 5. The tabulation begins with  $\varphi = 30^{\circ}$ , because at this value the incoming ray hits the vertex after reflection. At smaller values of  $\varphi$  the ray follows a polygon path before it hits the optical exis. No attempt has been made in this first survey to analyze this portion of the incoming radiation. Qualitatively, it increases the radiation influx in the outer portion of the focal line  $(0.7 \le r \le 1.0)$ .

The ratio f/r as well as eq/ds are plotted in Fig. 12. The curve f/r shows the well-known fact that for spherical reflectors of very small are ture a focal point exists which is located at 0.5 r. The curve dq/ds indicates that for large aperture spherical reflectors the maximum differential radiation flux is at  $\varphi = 45^{\circ}$ , corresponding to 0.7 r.

Using Eq. (1-11) the integral radiation flux can be computed. The result is shown in Figure 13, where Q is given in terms of  $<\pi$  r<sup>2</sup> Sens function of  $\varphi$ . In accordance with the trend shown by the dq/ds curves, the Q curves indicate that there is little or no contribution from radiation influent the extreme values of  $\varphi$ . This trend is the same as for d in (16) for the perabelic mirror.

C

By computing average values of q, Eq. (1-12), one can find the temperature of a black body receiver (flat) along  $0.5 \le r \le 1.0$ , by using the Stefan-Boltzmann relation  $T = (q/G^*)^{\frac{1}{4}}$  where  $G^*$  is the Stefan-Boltzmann constant. This has been done by using average values of q in intervals of five degrees. The result is shown in Fig. 14 for the radiation flux density per our of minimum area and for the temperature. This temperature, representing the highest theoretically attainable in a reflector of this type  $(R_n = 1.0)$ , is  $1,000^\circ \text{K}$  or  $1,000^\circ \text{K}$  lower than the theoretical maximum found in (16) for the parabolic mirror. The values of q are also presented in Fig. 10 where they are presented in the engineering system.

#### References

- 2ur Theorie der Photonenraketen, by Eugen Saenger, Fifth International Astronautical Congress, Zuerich, Switzerland, 1953, cf. Space Flight Problems, pp. 32-40, Laubscher & Cie Publ., Biel-Bienne, Switzerland, 1954.
- Wege Zur haumschiffahrt, by Hermann Oberth, R. Oldenbourg Publ., Huenchen and Berlin, 1929.
- The Atomic Rocket, by Leslie R. Shepherd and A. V. Cleaver, J. Brit. Interplanetary Soc., vol. 7, pp. 184, 234, 1948 and vol. 8, pp. 23, 59, 1949.
- Interplanetary Transport Technique, by H. Preston-Thomas, J. Brit. Interplanetary Sec. vol. 11. np.173 ff., 1952.
- A Comparison of Propellants and Working Fluids for Rocket Propulsion, by Krafft A. Ehricke, J. Am. Rocket Soc., vol. 23, no. 6, p. 287, 1953
- Facic Design Principles Applicable to Reaction-Provelled Space Vehicles, by Darrell C. Romick, Proceedings of the Fifth International Astronautical Congress, pp. 21-99, Innobruck, Austria, August 5-7, 1954, F. Hecht edit., Springer Publishing Co., Vienna, 1955.
- Possibilities of Electric Space Ship Propulsion, by Ernst Stuhlinger, same as (6), pp. 100-119.
- Electrical Propulsion for Space Ships: with Nuclear power Source, by Ernst Stuhlinger, The Journal of Astronautics, vol. 2, no. 4, Winter, 1955.

  The American Astronautical Society.
- 9 Stationaere Kernverbrennung in Baketen, by Eugen Saenger, Astronautica Acta, vol. 1, p. 61, 1955.
- 10 Electrical Wind Phenomena, by H. C. Early, H. L. Smith, and D. C. Lu, Su,, ary Report no. 1, Nov. 1952, Engineering Research Institute, Univ. of Michigan, Ann Arbor.
- 11. Test Facilities for Ultra-High-Speed Aerodynamics, by R. Smelt, Arnold Engineering Development Center Rep. AEDC-TR-55-6, Air Research and Development Command, USAF, June, 1955.
- 10 Liquid Hydrogen .... Field Use and New Liquefaction Systems, by Alexis Pastuhov and C. Lincoln Jowett, Arthur D. Little Inc., Cambridge, Mass.,
- Taschenbuch fuer Chemiker und Physiker, by J. D'Ans and E. Iax edit., Berlin: Springer Publ., 1943.
- 14 Evaporation Nate of Liquid Welium, I, by Aaron Wexler, J. Appl. Physics, vol. 22, no. 14, pp. 1463-1470, Dec. 1951.
- Gomoilation of Thermal Proporties of Mydrogen in its Various Isotopic and Ortho-Para-Modifications, by M. W. Woolley, R. B. Scott and F. G. Brickwedde, Nat. Bur. Standards Ros. Paper RP 1932, vol. 41. Nov. 1941.

- Analysis of Large Aperture Parabolic reflectors for Solar Furnaces, by J. Farber and B. I. Davis, Convair, A Division of General Dynamics, Rep. TRM-77, Jan. 1954, San Diego, Calif.
- Du Pont 'vlar Polyester Film, Physical, Electrical and Chemical Properties, Technical Bulletin No. 1-2-53, Film Department, E. I. Du Pont De Nemours & Co. (Inc.), Wilmington, Del.

3

TARE 1 SOFT CHRESTICS OF PROPILETON SYSTEMS

Character for the con-	Chemical Combustion	PIJe Heating	Soler-Heating	Ast the time	Electric (Tops)
				Soler or	
Secuence Conversion	Chemical Thermal	Muclear	Soler	Pile or 6 - Desay	Solar or
. #	Kimtle	Kinetie	Kinetic		<i>x</i> =
Rance of Epecific Impulses (Ib sec/lb)	200-300 (surface) 300-400 (space)	606-900 (Tc= 2,000°K)	(705-500 (705-100°K)	700-1,200 (Depending on Design)	Kirtia 211,000 (Dependire on Design)
Systems Specific Thrust (1b thrust/1b purp, 17st, hardware)	50 - 80 (Engline, Pump, Pressur- isation Thrust Structure Auxiliary System)	Probably less but same order of merri- tude (without shield- ine), Depends on engine	50.05	0.005-0.0015 (\$4- Pacay) 0.005-0.0015 (Solar) (Perending on Design)	~0.007
Specific Energy Computation (kv/lb thrust)	Syst.) d to best n Jet)	15-20 (Pile)  — Lang (Corrected to to heat of medius)  10-15 (Corrected in let)	12-14 (Propolation Syri) 11-13 (Bared on Mirror 13se R = 0, 9-11 (Corrected to heat of medium) Fig. (Corrected to 3et)	60-97 (Prepaleion Syst) 20-30 (Converted to bests of medium) 14-21 (Converted in Jet)	~300 (Corrected in Jet.)
Acceleration (Thrust/at)	1.2-1.5 (Surface) ≥ 0.25 (Space)	1	0-2	5.10-4 - 5.10-3	10-5 - 5.10-5
Forting Fluid Fedius)	Puel & Oxidiser	Mi 3+ He+ H2	2,4	Producting medium H2 CH2. T2	Casium, Rubidium or other lons
Maneuverability	Excellent	Excellent	Poor	Poor	Users Days
Areas of Operation	Space (Moon, Mars, Verms) Space	(Verus-Care)	Cutside Shadow)	Cislumar and Lunar Space	Interplantary Space (Inner and outer
Probability of Extra- ferrestrial Supply of Medium,	Horse	Volume and Possibly outer solar system (Titan)	Note: Serve	Libral Solat System	solar system) The best media (lowest ionisation potential) not readily available extra-
fereral fonciousions:  (a) Initial investment in spaceborne hardware in the figure of the ferenced	vol	Padriy Lov	Fairly High	414	Very Hich
(b) Tunning Lupply in work-  for Fluid (Medium)  for Frould don Switce Wain- Very High  for France & Feliability  actively company	Very High Very Inited life of engine & Umps: compar- atively committee	Pairly High Long life, Removal of firsion products necessary	Pairly High Lour life; simple syst. Maint, depends on density of metencials	Low Carnot be judged reliably at present	Very Low Long life; complex system
ture Circlistre Areas of Long of Englished.	Ascent, Landing (aktiese)	t. Landing (air- Lungr-Varra-Venus for	olar syrtem (?)	Cislurar Inner solar system (2)	Interplanetary
		oupties of the second			i

Table 2 Comparison of Hydrogen and Halium

	Hydrocen	Helium
Molecular Weight	2.016	4.003
Freezing Point	-436°F (12) -260°C	-459.4°P -273°C (13)
Boiling Point (760mm)	-423°P(12) -252.8°C	
Sp. Gravity (liq.,760mm)	0.071 (4.43 lb/rt <sup>3</sup> )(13)	0.13 (13) (8.1 1b/ft <sup>3</sup> )
Critical Pressure	13.5 atm (12) 12.8 atm (13)	2.28 atm (13)
Critical Temperature	-400°F (12) -240°C	
Critical Sp.Cravity	0.031 (13) (1.94 lb/ft³)	0.069 (13) (4.3 lb/rt3)
Specific Heat (Average between B.P. and 3000°k)	3.7 Btu/lb°F 3.7 cel/g°k(5	1.13 Btu/lb°F ) 1.13 cal/g°k (5)
Heat of Vaporization (760mm)	108 cal/g 194 Btu/lb (12	4.93 cal/g (4.93°k) (14) 2) 8.87 Btu/lb

# Table 3 Some Relevant Properties of the Polyester MYLAR (17)

Molting Point

Specific Gravity

Thermal Conductivity

Thermal Coefficient of Linear Expansion

Light Transmission

Tensile Strength

Bursting Strength

Oxygen Permeability

Hydrogen Permeability

250-255°C (about 490°F)

1.38-1.39

3.63.10-4 cal/cm sec °C

20.10-6 °F

about 90% in visible (>4,000 Å)

zero at <3,000 Å

25,000-30,000 pai (about 0°F) 10,000 pai @ 300°F (150°C)

45 lb @ 1 mil thickness

0.9 g/100 m<sup>2</sup> hr 4 1 mil thickness

No Data Available

# Table 4 Characteristic Data of the Solar Powered Spaceship Prototype

1. Weights		
Madiation Collectors (2)		1,000 16.
Polyester Spheres (2)	740 lb.	1, 00 10.
Cold Tubes, Vires, Springs	190	
Connections.Reflector Drive, Misc.	70	
Not Piping System (for 2 Collectors)	• •	700
✓ Heating Elements	200-1ъ	<i>,.,</i> ,
	350	
Hot Pining		
Padiation Jacket for Hot Pipes	150	500
The ine	(0.15	500
Turbine and Alternator	60 lb.	
Low-Pressure Booster Pump (0.37 1b/sec Hg)	10	
Wi-Pressure Pump & Motor (2 sets)	50	
Miring (2 sets) (electrical)	24)	
Motor	50	
Connections & Shut-off Valves	10	
Array of Colid Prop. Starter Rockets	300	
Liquid Hydrogen Tank (17 ft Dia., 900 ft Surface)		۶,00
Centola, Crew & Equipment		2,400
Liquid Hydrogen		11,000
Gross Weight		16,400 1b
Dry Weight		5,600 lb
2. Miscellaneous Data		
Collector Sphere: Diameter		120 ft.
Intercented Area, A.		12.370 ft.4
Circumference		402.1 ft. <sub>2</sub>
Surface Area		51,468.8 ft <sup>2</sup>
Volume		1,098,000 ft <sup>3</sup>
Wydrogen gas wt. in sphere (o.ol psi) 300°F		2.74 lb
-150°F		6.95
Wellium gas wt. is roughly twice that of H2		0.37
Intercented area theoretically required to produce 30	Th. of	
thrust at $T_{cn} = 450$ sec		9,300 ft <sup>2</sup>
Reflector efficiency 9,300/12,370		0.725
Theoretically produced thrust per reflector		111 1ь
Therew theoretically collected by reflector		1,287 kw
Theoretical specific energy consumption		
Actual thrust assumed to be produced per collector		12.9 kw/lb thrust
Total thrust produced Thrust-to-weight ratio: initial		160 1b
- 1 <del>- 1 - 1 - 1 - 1 - 1 - 1 - 1 - 1 - 1</del>		2 4 7 4 7 4 7 5
maximum final		2.963.10-2
3. Ideal Performance		
Touting Factor (1.49/2ross Mt.)		2.67
Waximum Possible Mass Satio Based on Loading Factor		3.03
Operational Constitio Impulse Assumed		450 lb sec/lb
Limit Valority Limit on Above Data		15,730 ft/sec
The state of the s		- 79 - 70 - 7 0 - 110 C

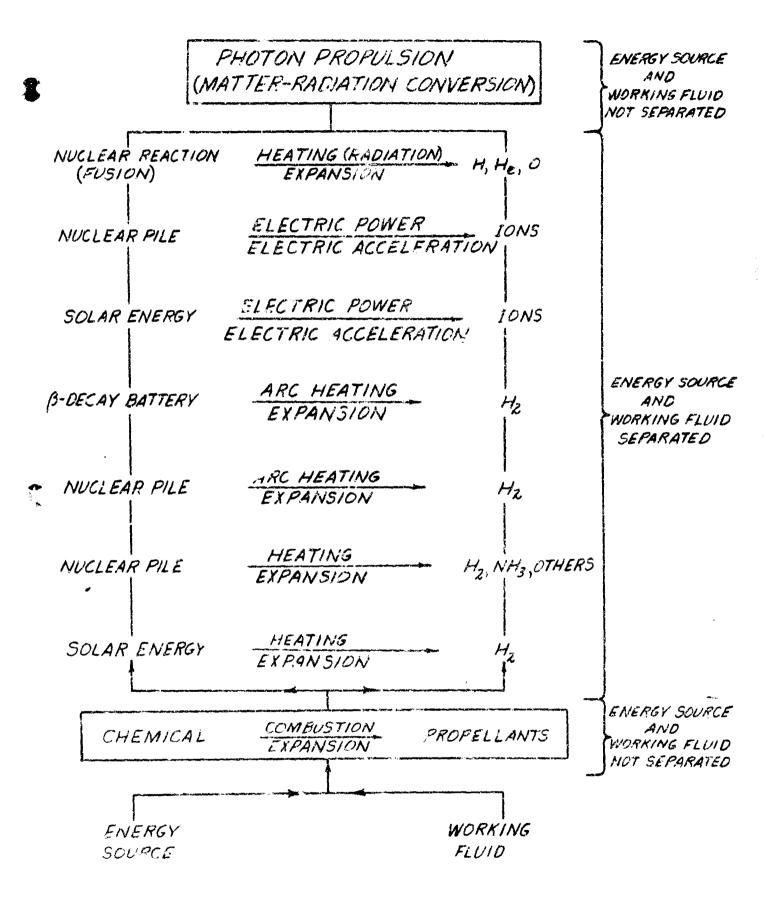
The level performance characteristics indicate that this prototyre is not sended on a simulation flight and return into the orbit of departure, taken at 100 mail. Third. I compare of the prototype is to study problems of design, and month to be the control of design of the vehicle.

Table 5 Characteristic Functions of the Spherical Reflector

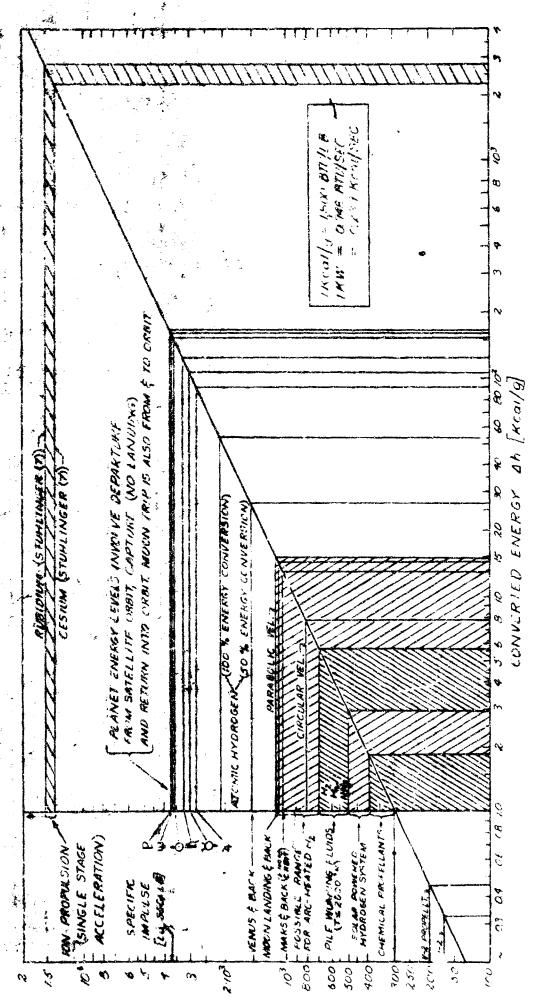
<b>∳</b> (deg)	( - )	t <sub>1</sub> (p)	f <sub>2</sub> (φ)	£3( Y)	
30	1.0	0.3248	0.03541	0.07812	
35	0.8717	0.4149	0.05048	0.12675	
40	0 <b>.777</b> 8	0.4775	0.06658	0.18554	
45	9 <b>.7071</b>	0.5000	0.08250	0.25000	
50	0.6527	0.4775	<b>0.0970</b> 8	0.31441	
55	0.6104	0.4149	0.10945	0.37325	
60	0.5773	0.3248	0.11907	0.42187	
65	0.5517	0.2248	0.12585	0.45785	
70	0.5321	0.1327	0.13005	0.48107	
75	0.5176	0.0625	0.13224	0.49358	
80	0.5077	0.0200	0.13310	0.49866	
€5	0.5019	0.00261	0.13332	0.49990	
89	0.500				
90	-	0	0.13339	0.50000	

I

Z: 5



FIGI CORRELATION BETWEEN ENERGY SOURCE AND WORKING FLUID IN VARIOUS PROPULSION SYSTEMS



The later was to be a few and the was the second

FIG & ENERGY SPECTRUM OF PROPULSION SYSTEMS AND FLIGHT PERFORMANCE

The second of th

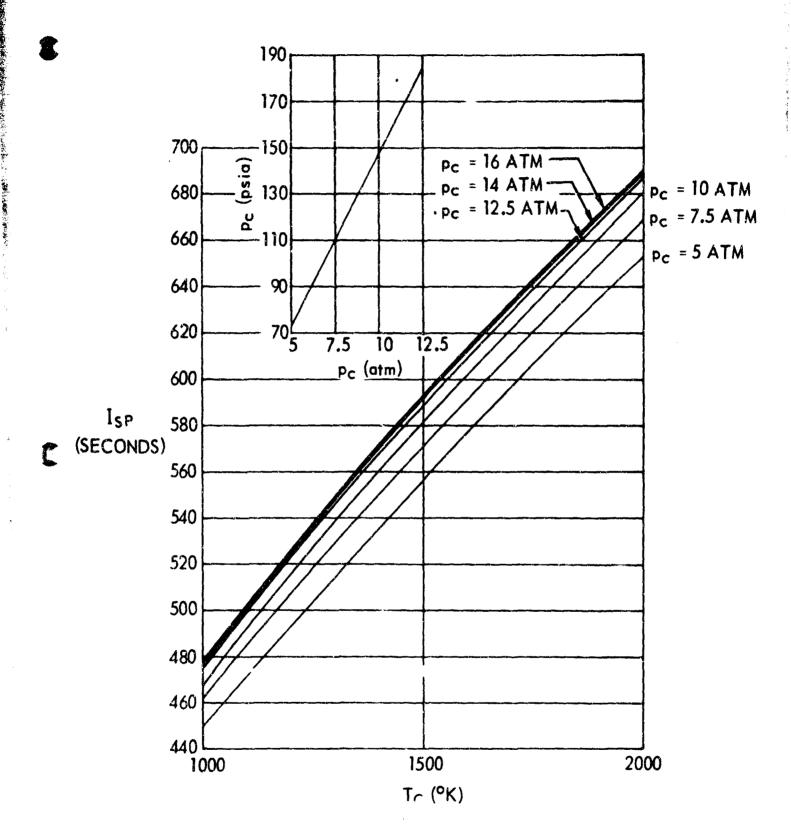
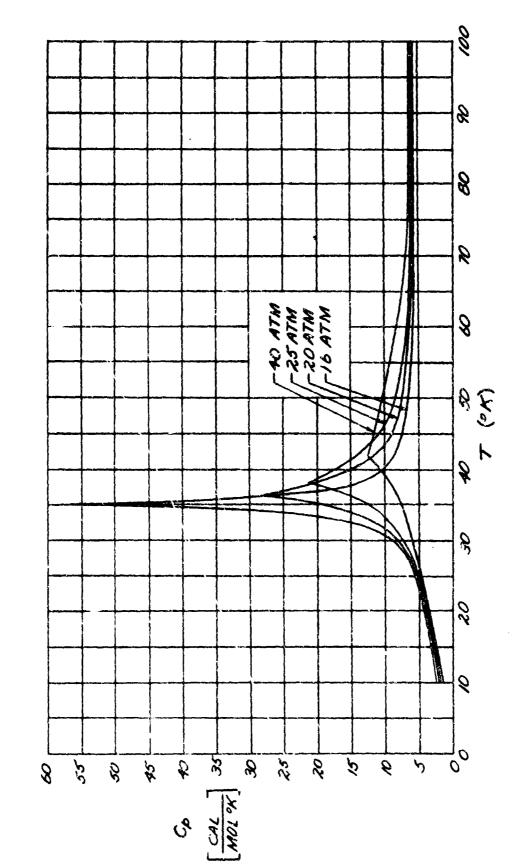


FIGURE 4. SPECIFIC IMPULSE OF PURE HYDROGEN AS FUNCTION OF CHAMBER TEMPERATURE

MOLECULAR HEAT OF LIQUID HYDROGEN AT SUPERCRITICAL PRESSURE F16. 5



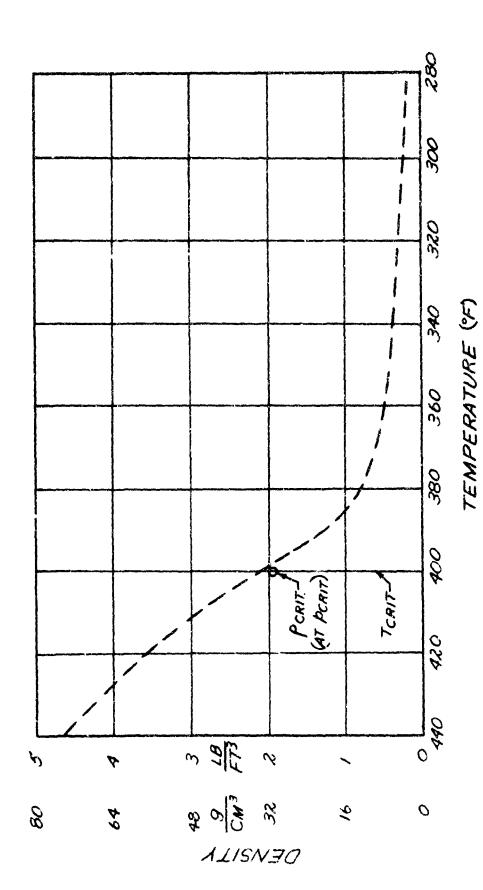
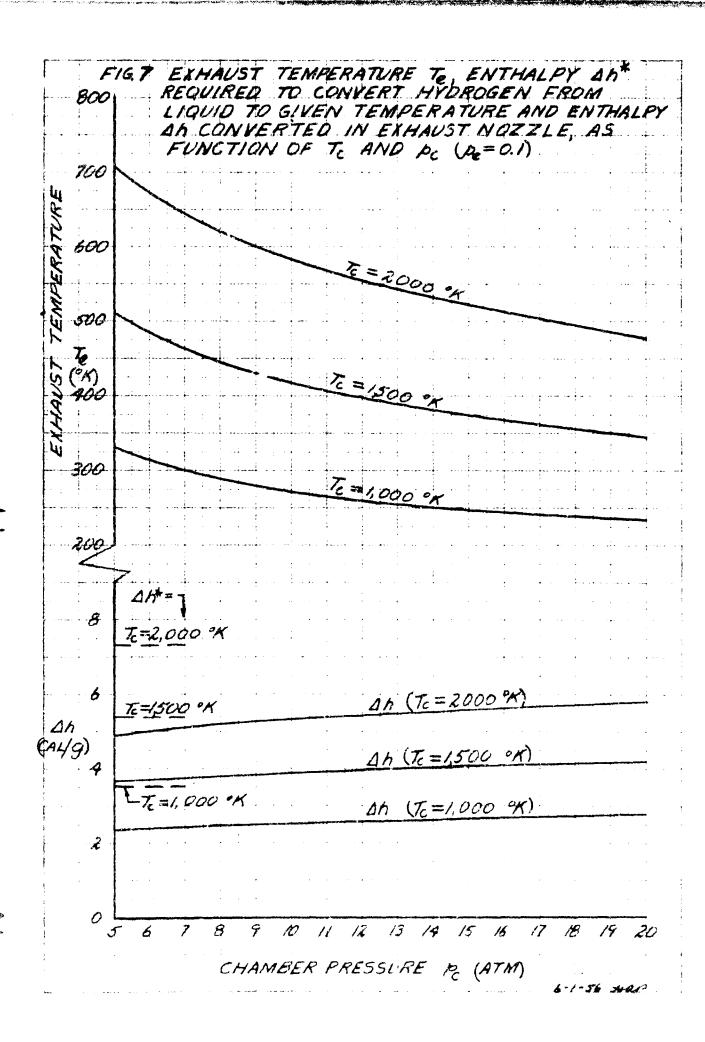


FIG & ESTIMATED DENSITY OF HYDROGEN IN SUPERCRITICAL STATE ((SATM)

C



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Action of the board of the con-

FIGURE // SPERICAL REFIECTOR

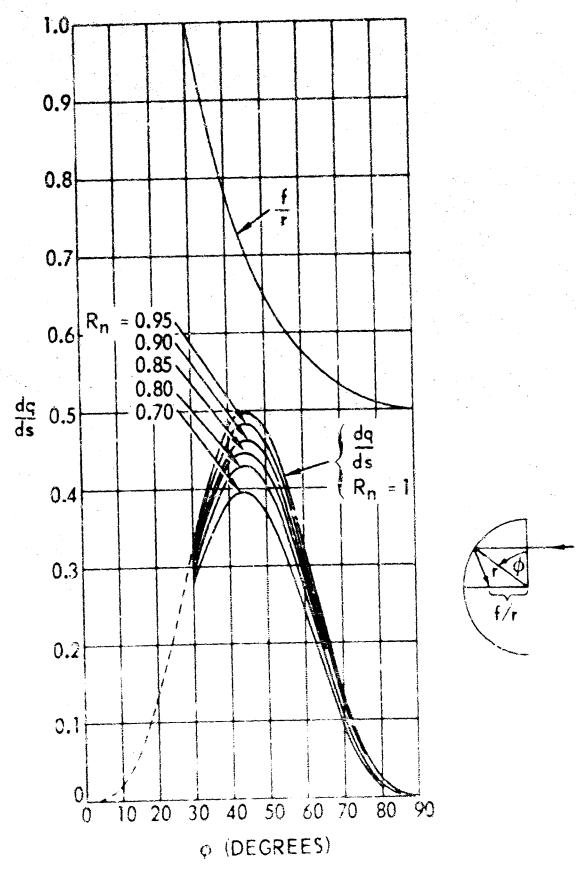
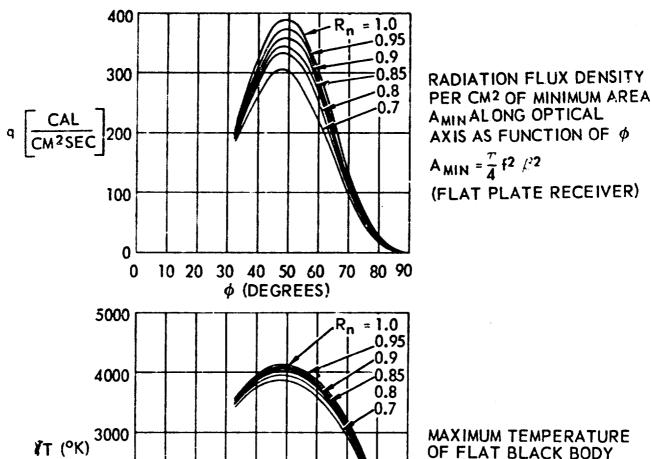


FIGURE 12. HEMISPHERICAL MIRROR:
DISTANCE OF LUCAL FOCAL POINT FROM CENTER
OF SPHERE & DIFFERENTIAL RADIATION FLUX dg
ds

a os Q-66 004 608 20 30 70 80 R f (15055ES) !



2000

1000

0

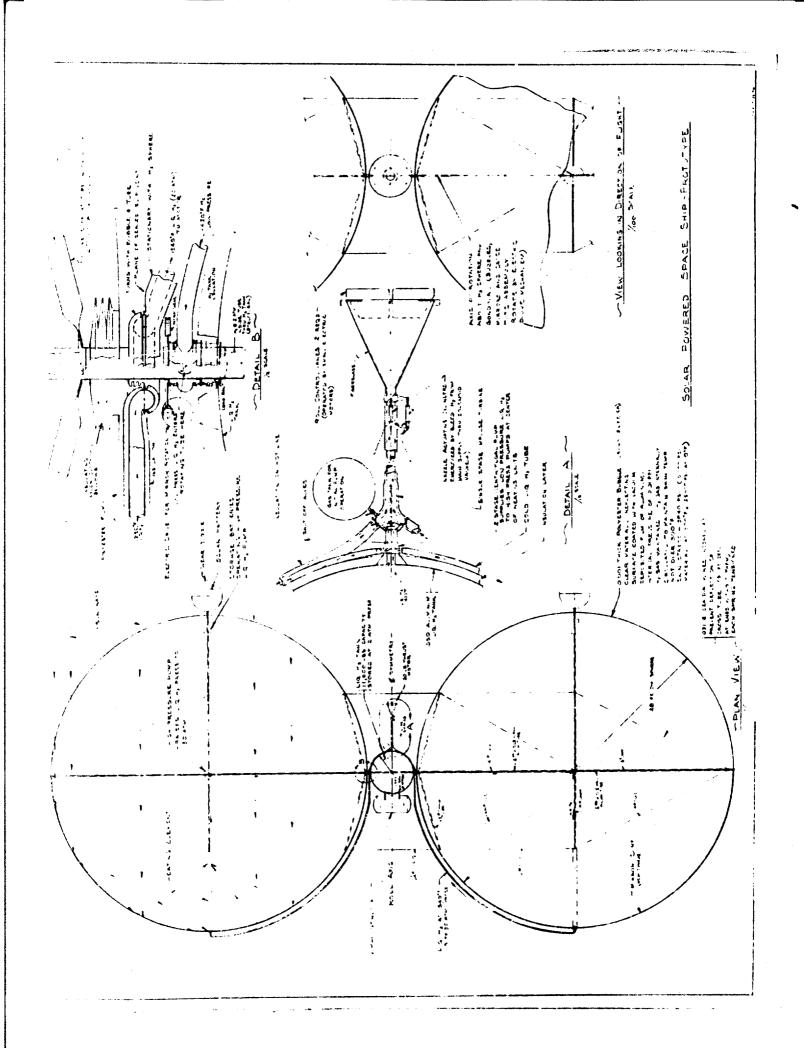
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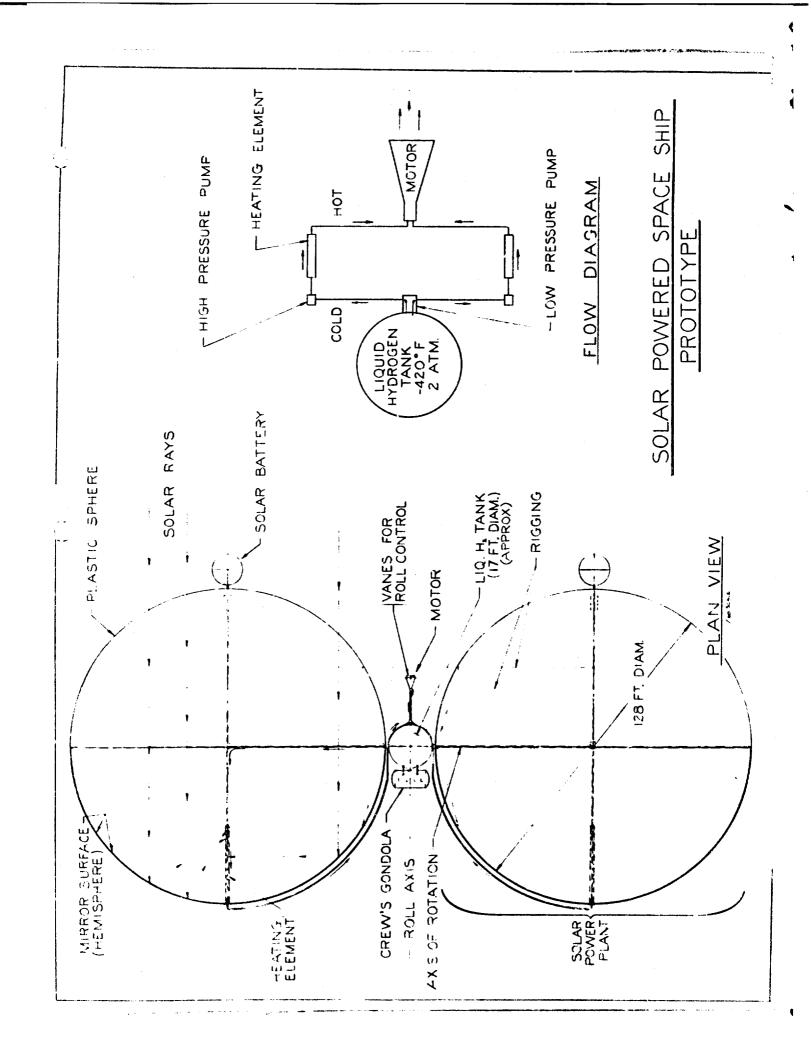
MAXIMUM TEMPERATURE OF FLAT BLACK BODY RECEIVER IN OPTICAL AXIS AS FUNCTION OF  $\phi$ 

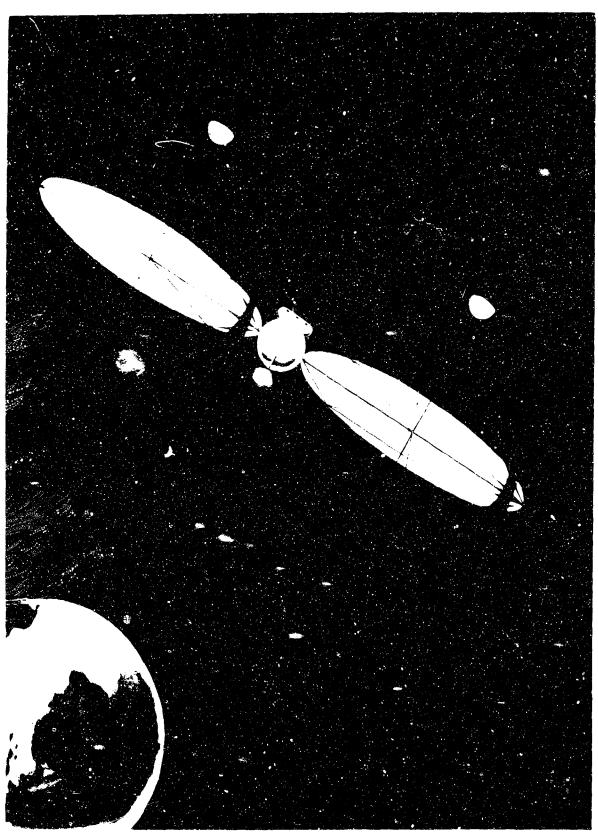
FIGURE 14. ENERGY CHARACTERISTICS OF SPHERICAL MIRROR (IN SPACE)

10 20 30 40 50 60 70 80 90

φ (DEGREES)







The Solah Powered Space Ship